

## Radio Astronomy Advanced Course II: Report Assignment

Answer the following problems and send you report to [takeshi.okuda@nao.ac.jp](mailto:takeshi.okuda@nao.ac.jp) by email with “Radio Astronomy Report 2019” in the subject line. Its due date is August 9<sup>th</sup> (Fri), 2019 (JST). The report shall be written in English or Japanese as a PDF file within 10 MB. A scan of your report written by hand is also acceptable. I will send you the acknowledge message within three days after you send it. Please contact me again by email and add [takeshi.okuda@alma.cl](mailto:takeshi.okuda@alma.cl) in CC if you don’t receive the acknowledgement message from me. Enjoy!

July 19<sup>th</sup>, 2019  
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1.

- A) Suppose that a distribution of brightness temperature of a planet  $T_B(\theta, \phi)$  and a normalized beam pattern  $P_n(\theta, \phi)$  of a radio telescope are given by,

$$T_B(\theta, \phi) = \begin{cases} T_{B0} & \text{for } \theta \leq \theta_p \\ 0 & \text{for } \theta > \theta_p \end{cases}$$
$$P_n(\theta, \phi) = \exp \left[ -4 \ln 2 \left( \frac{\theta}{\theta_b} \right)^2 \right]$$

where  $\theta_p$  and  $\theta_b$  are an angular diameter of the planet and FWHM of a beam, respectively. Obtain an antenna temperature  $T_A^*(\theta, \phi)$  and an aperture efficiency  $\eta_A$  when observing the planet.

- B) The telescope has a diameter of 10 m and a beam size ( $\theta_b$ ) of 21 arcsec at 350 GHz, and brightness temperature ( $T_{B0}$ ) and an angular diameter ( $\theta_p$ ) of the planet are 130 K and 6 arcsec, respectively. If  $T_A^*$  is 6.0 K, obtain an aperture efficiency and a main beam efficiency of the telescope.
- C) If  $\eta_A = 0.8\eta_{sf}$  ( $\eta_{sf}$ : surface efficiency), obtain surface accuracy of the telescope.

2.

- A) Suppose that you will observe towards one point of an object by CO(J=1-0) by a position switched observation with a single dish telescope (FWHM=20 arcsec). If the system temperature, a frequency bandwidth of a spectrometer, integration time, and quantization levels in the spectrometer are 200 K, 1 MHz, 60 seconds, and 8-levels, estimate achieved  $\Delta T_{\text{rms}}$ . Use an equation for the sensitivity estimation in Lecture 5.

B) The systemic velocity of the object and a rest frequency of CO(J=1-0) are 0 km/s and 115.271 GHz, respectively. Suppose that  $X_{\text{CO}} = 3.0 \times 10^{20} \text{ [cm}^{-2} \text{ (K km s}^{-1}\text{)}^{-1}\text{]}$  in the object and a distance to the object is 5.0 Mpc, obtain 1  $\sigma$  upper limit of molecular hydrogen mass in the observation when integrating over 200 km/s.